

INSTITUT FÜR ISOTOPENFORSCHUNG UND KERNPHYSIK
DER UNIVERSITÄT WIEN

E I N L A D U N G

zum

I N S T I T U T S S E M I N A R

von

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**A thick-target measurement of the cross section
of the supernova-nucleosynthesis $^{40}\text{Ca}(\alpha,\gamma)^{44}\text{Ti}$
reaction by accelerator mass spectrometry**

The ^{44}Ti ($t_{1/2} = 59$ y) nuclide is considered an important signature of core-collapse supernova nucleosynthesis. ^{44}Ti has recently been observed as live radioactivity by γ -ray astronomy from the Cas A supernova remnant, showing yields larger than expected from current stellar calculations. We investigated in the laboratory the major ^{44}Ti production reaction, $^{40}\text{Ca}(\alpha,\gamma)^{44}\text{Ti}$ ($E_{\text{cm}} \sim 0.6\text{-}1.2$ MeV/u) by off-line counting of ^{44}Ti nuclei using accelerator mass spectrometry. The observed yield is significantly higher than inferred from previous prompt γ -spectroscopy experiments and is well reproduced by a microscopic statistical model calculation. The derived astrophysical rate of the $^{40}\text{Ca}(\alpha,\gamma)^{44}\text{Ti}$ reaction is a factor 5-10 higher than calculated in current stellar models and results in an increase of the calculated SN ^{44}Ti yield by a factor ~ 2 over current estimates.

Dienstag, 29. August 2006, 11:00 Uhr s.t.

**1090 Wien, Währinger Str. 17, "Kavalierstrakt",
1. Stock, Seminarraum von VERA**

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W. Kutschera