## <sup>10</sup>Be CONTENT IN SUEVITE BRECCIA CLASTS FROM THE BOSUMTWI CRATER FILL AS A PROXY FOR THE CONTENT OF SURFACE COMPONENTS.

A. Losiak<sup>1</sup>, E. Wild<sup>2</sup>, L. Michlmayr<sup>2</sup> and C. Koeberl<sup>1,3</sup>. <sup>1</sup>Dept. Lithospheric Research, Univ. Vienna, 1090 Vienna, Austria (anna.losiak@univie.ac.at). <sup>2</sup> VERA Laboratory, Faculty of Physics, Isotope Research, Univ. Vienna, 1090 Vienna, Austria. <sup>3</sup>Natural History Museum, 1010 Vienna, Austria.

**Introduction:** According to the current understanding of meteorite impact processes, surface target material is transported from a crater in the form of ejecta or is vaporized/melted (e.g., [1]). The formation model of tektites from the surface of the target rocks has been established using the <sup>10</sup>Be content of tektites (e.g., [2]), and chemical comparison with the possible target surface material (e.g., [3]); it was also reproduced by computer modeling (e.g., [4]). On the other hand, some observations ([5, 6]) suggest that part of the surface material may be incorporated into the crater-fill. The aim of this study is to check if surfacederived material is present in suevitic breccias to better understand formation mechanisms of fallback breccias.

**Samples:** The Bosumtwi crater was chosen as study site because of its relatively large size (10.5 km in diameter), relatively young age of 1.07 Ma [7], good preservation, and availability of core samples. Clasts from suevitic breccia selected for this study come from the LB-07A and LB-08A cores that are located within the crater and represent fallback breccia (e.g., [7]). Of 18 analyzed clast samples, 13 came from core LB-07A and represent depths of 336.4 – 405.7 m and 5 are from core LB-08A from depths 239.5 – 264.9 m.

**Methods:** For each sample, 0.8 g of finely grounded material from clasts containing *in situ* produced and meteoric <sup>10</sup>Be was dissolved in a mixture of HF and HNO<sub>3</sub> by microwave digestion. A carrier (1 mg, <sup>10</sup>Be/<sup>9</sup>Be ratio:  $2.82\pm0.31*10^{-15}$  [2 $\sigma$  uncertainty]) was added to the sample, and then Be was chemically separated from the sample solution. <sup>10</sup>Be/<sup>9</sup>Be ratios were measured at the Vienna Environmental Research Accelerator Facility (VERA) at the University of Vienna.

**Results:** Most samples have  ${}^{10}\text{Be}/{}^9\text{Be}$  ratios indistinguishable from the blank within  $2\sigma$  uncertainty. However, one sample has a significantly higher  ${}^{10}\text{Be}/{}^9\text{Be}$  ratio (7.94±0.84\*10<sup>-15</sup>[2 $\sigma$ ]).

**Discussion:** The data obtained so far suggest that none of the analyzed clasts present in the suevitic breccia of the Bosumtwi crater come from depths less than 15 meters, or/and are contaminated with meteoric <sup>10</sup>Be carried in the rain-water. The reasons for one of the samples having an elevated <sup>10</sup>Be content could be: 1) the clast originated from a depth of ~20-25 meters below the surface level within the impact crater target material, 2) the clast was in contact with water enriched in meteoric <sup>10</sup>Be, 3) experimental uncertainty (to exclude this possibility, the measurement will be repeated).

**Acknowledgment:** Supported by University of Vienna doctoral school IK-1045 (Planetology).

**References:** [1] Melosh H.J. 1988. Impact Cratering. [2] Serefiddin F. et al. 2007. Geochimica et Cosmochimica Acta 71:1574-1582. [3] Son T.H. and Koeberl C. 2007. GFF 29:161-176. [4] Artemieva N.A. 2000. In: Impacts in Precambrian Shields, Springer, pp. 257-276. [5] Puura V. et al. 2004. Meteoritics & Planetary Science 39:425–451. [6] Reimold W.U., et al. 1992. Geology 20:1079-1082. [7] Koeberl C. et al. 2007. Meteoritics & Planetary Science 42:483–511.